

Cosmological Constant, Vacuum Fluctuations & the Equivalence Principle

EMILIO ELIZALDE

ICE/CSIC & IEEC, UAB, Barcelona

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Outline of this presentation

- Einstein's Cosmological Constant

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- Gravity Eqs as Eqs of State

Einstein's Cosmological Constant

Our universe seems to be spatially flat and to possess a non-vanishing cosmological constant

- For cosmologists and general relativists: a great mistake why was it put there, in the first place? (Einstein)
- For elementary particle physicists: a great embarrassment no way to get rid off (Coleman, Weinberg, Polchinski)
- The cc Λ is indeed a peculiar quantity
 - has to do with cosmology Einstein's eqs., FRW universe
 - has to do with the local structure of elementary particle physics stress-energy density μ of the vacuum

$$L_{cc} = \int d^4x \sqrt{-g} \mu^4 = \frac{1}{8\pi G} \int d^4x \sqrt{-g} \lambda$$

In other words: two contributions on the same footing (Zel'dovich, 68)

$$\frac{\Lambda c^2}{8\pi G} + \frac{1}{\text{Vol}} \frac{\hbar c}{2} \sum_i \omega_i$$

Einstein Eqs, FLRW Sol, Hubble Const

Einstein Equations (1915-17): $G_{\mu\nu} - \lambda g_{\mu\nu} = -(8\pi G/c^4)T_{\mu\nu}$

Geometry = Energy-Matter

$G_{\mu\nu}$ linear combination of the metric $g_{\mu\nu}$ and 1st & 2nd derivatives

$T_{\mu\nu}$ energy-momentum tensor

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R = -(8\pi G/c^4)T_{\mu\nu} + \lambda g_{\mu\nu}$$

Schwarzschild solution (1916)

r, θ, φ comoving co

$$ds^2 = \left(1 - \frac{2MG}{r}\right) dt^2 - \left(1 - \frac{2MG}{r}\right)^{-1} dr^2 - r^2 d\theta^2 - r^2 \sin^2 \theta d\varphi^2$$

Friedmann-Lemaître-Robertson-Walker (1935-36) sol (A. Friedmann 1922)

$$ds^2 = dt^2 - a^2(t) \left(\frac{dr^2}{1 - kr^2} + r^2 d\theta^2 + r^2 \sin^2 \theta d\varphi^2 \right)$$

most general family: *homogeneous* and *isotropic*, k parameter $\pm 1, 0$

One field eq looks like Newtonian eq for the gravit pot: $\nabla^2 \phi = 4\pi G (\rho + 3p/c^2)$

density & pressure contribute to the gravit pot $\lambda = 8\pi G \rho_{vac}$, $p_{vac} = -\rho_{vac} c^2$

From the FRW metric and Einstein Eqs, an “equation of motion” of the universe

$$H^2 \equiv \left(\frac{\dot{a}}{a}\right)^2 = \frac{8\pi G}{3}\rho + \frac{\lambda}{3} - \frac{k}{a^2}$$

From GR to Cosmology

With the definitions:

$$\Omega_m \equiv \frac{8\pi G}{3H^2} \rho_m, \quad \Omega_\lambda \equiv \frac{\lambda}{3H^2}, \quad \Omega_k \equiv -\frac{k}{H^2}$$

The equation of motion becomes

$$\left(\frac{da}{dt}\right)^2 = H_0^2 \left[\frac{\Omega_m^{(0)}}{a} + a^2 \Omega_\lambda^{(0)} + \Omega_k^{(0)} \right]$$

(the superscript (o) represents quantities measured at the present time)

In other terms, **Friedmann equation in Cosmology:**

$$\frac{\dot{a}^2}{a^2} = H_0^2 \left[\Omega_R \left(\frac{a_0}{a}\right)^4 + \Omega_{NR} \left(\frac{a_0}{a}\right)^3 + \Omega_k \left(\frac{a_0}{a}\right)^2 + \Omega_\lambda \right]$$

Ω_R relativistic matter ($p_R = \frac{1}{3}\rho_R$; $\rho_R \propto a^{-4}$)

Mach's princ

Ω_{NR} nonrelativistic matter ($p_{NR} = 0$; $\rho_{NR} \propto a^{-3}$)

Ω_λ cosmological constant ($p_\lambda = -\rho_\lambda$; $\rho_\lambda = \text{const}$)

$\Omega = \Omega_R + \Omega_{NR} + \Omega_\lambda$ total energy density (cosmic triangle)

Zero point energy

QFT vacuum to vacuum transition: $\langle 0|H|0\rangle$

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Spectrum, normal ordering (harm oscill):

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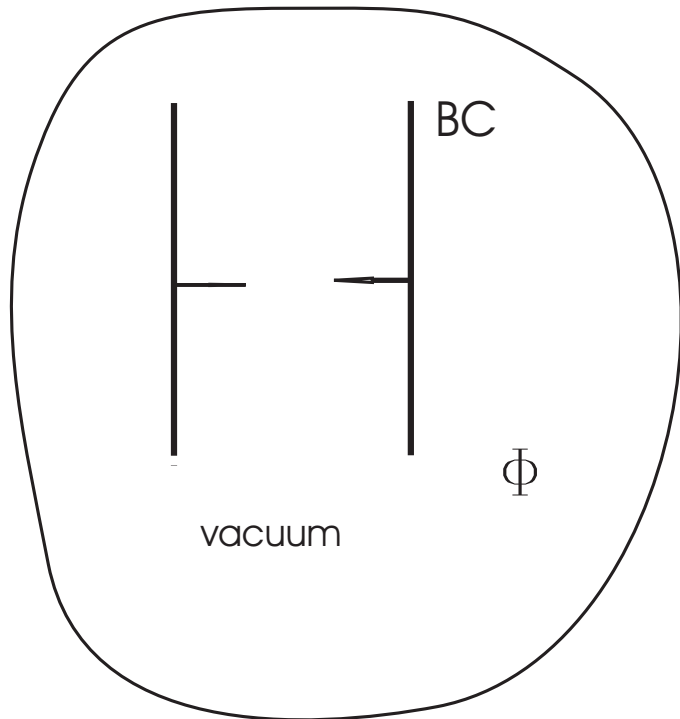
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Even then: Has the final value real sense ?

The Casimir Effect

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BC e.g. periodic

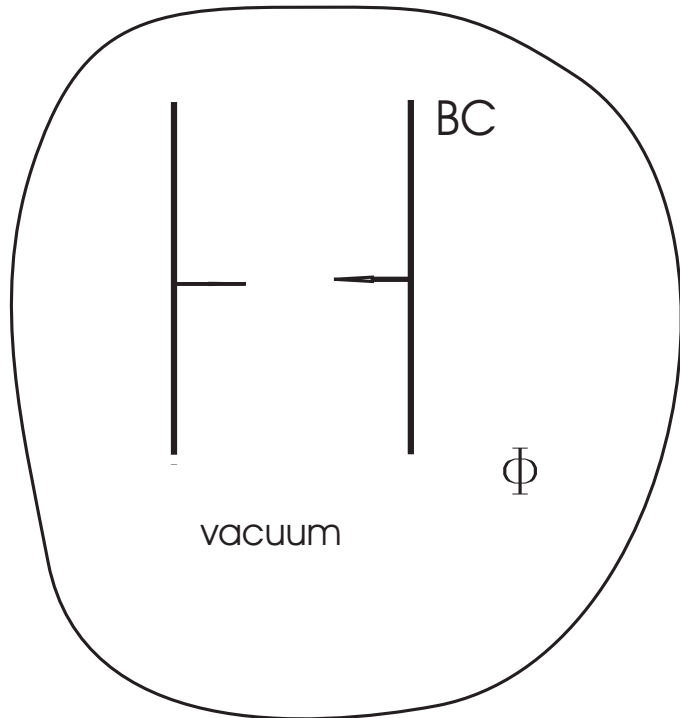


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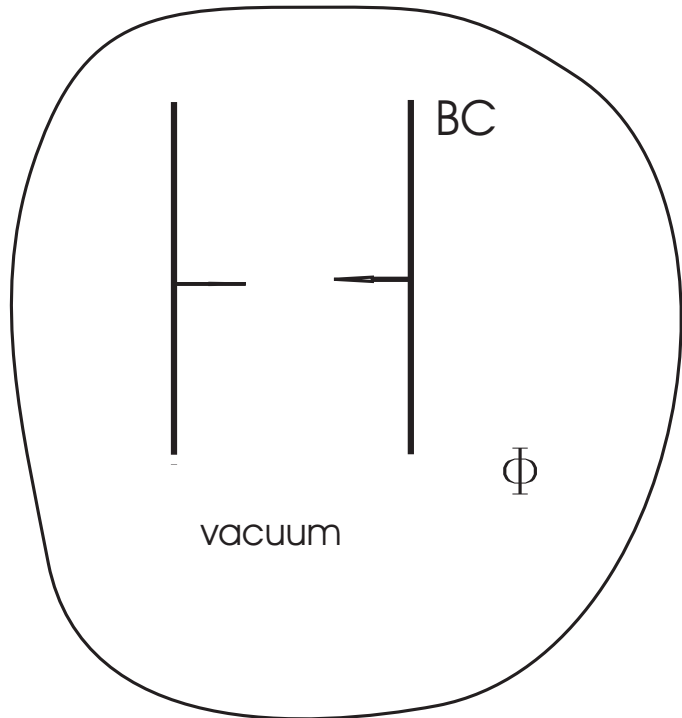
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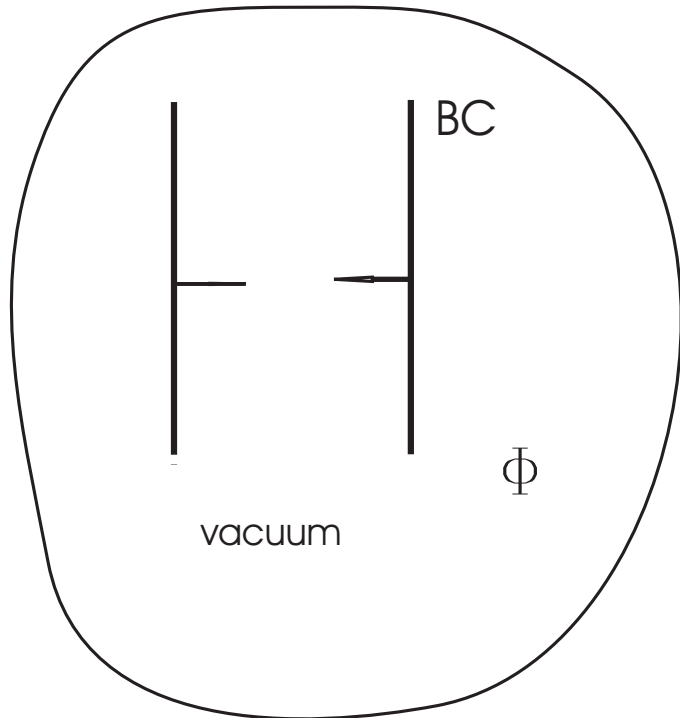
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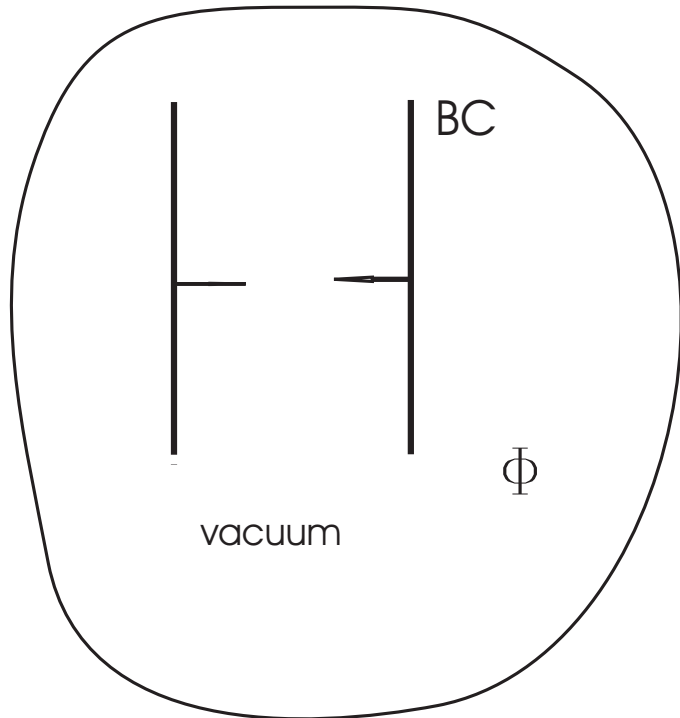
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Universal process:

The Casimir Effect



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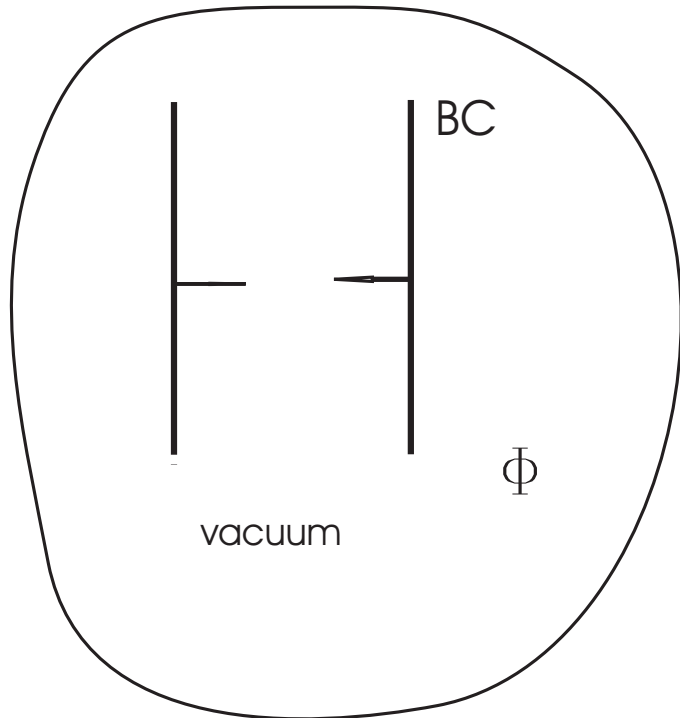
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- Sonoluminescence (Schwinger)
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- Optical cavities
- Direct experim. confirmation

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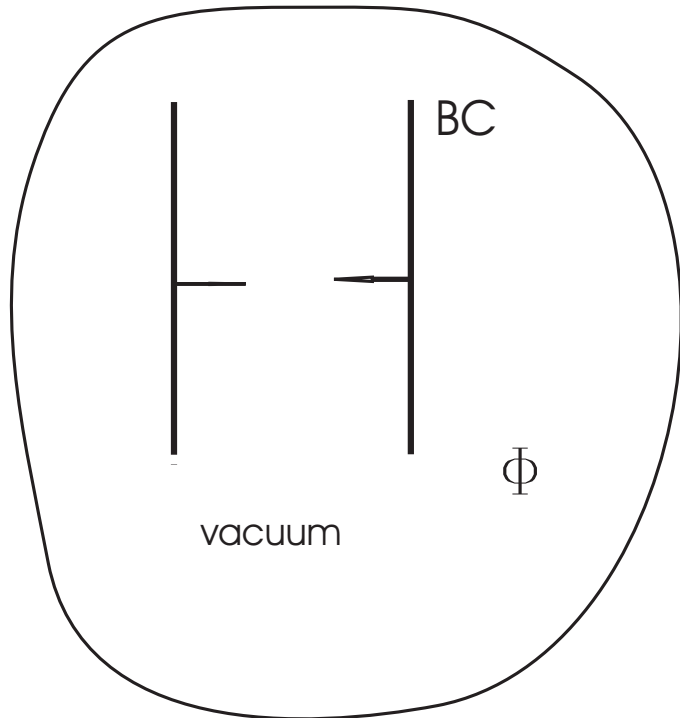
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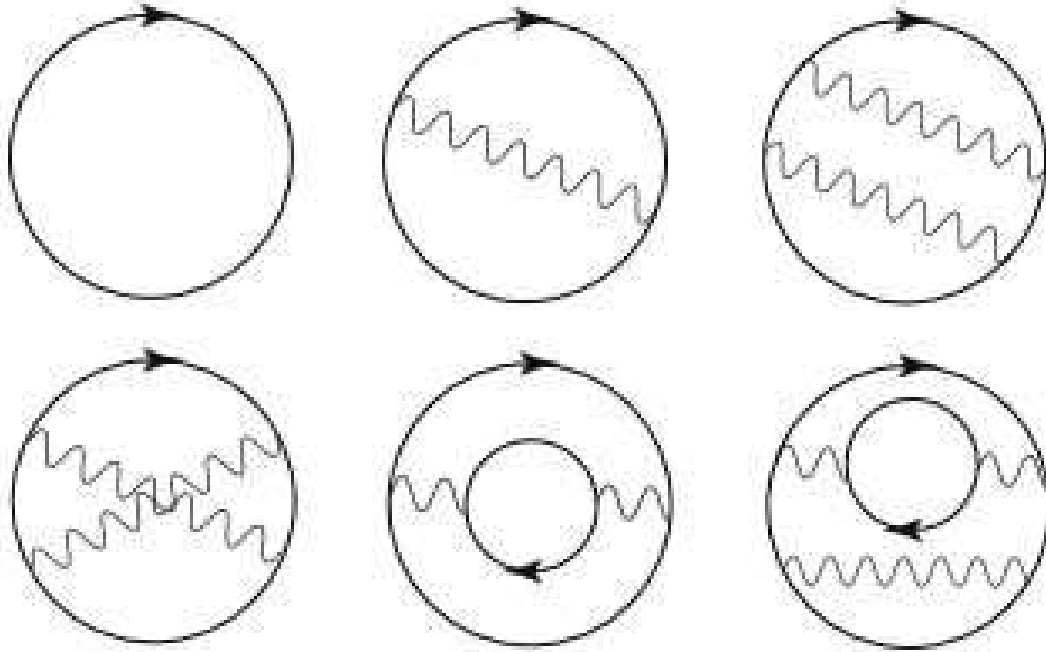
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- Dynamical CE \Leftarrow
- Lateral CE
- Extract energy from vacuum
- CE and the cosmological constant \Leftarrow

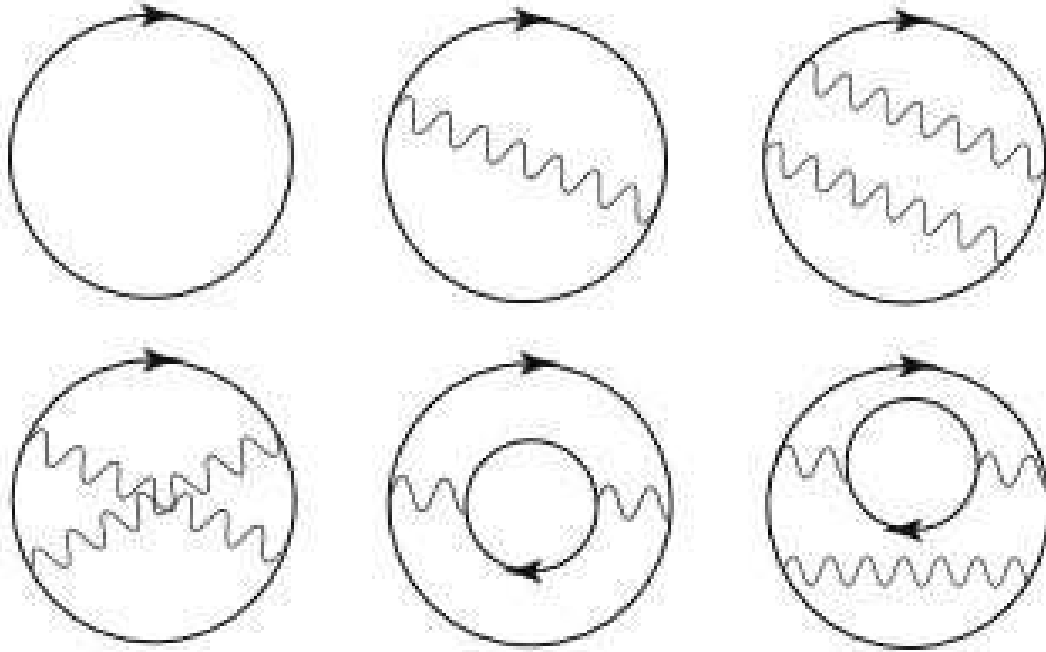
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⇒ Casimir force: calculated by computing change in zero point energy of the em field

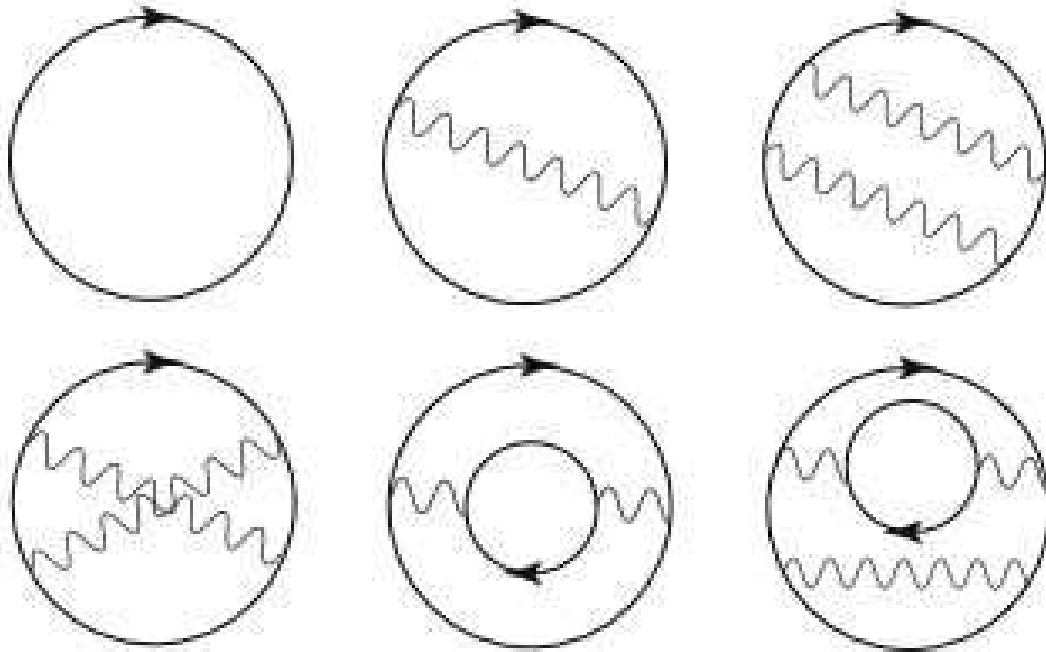
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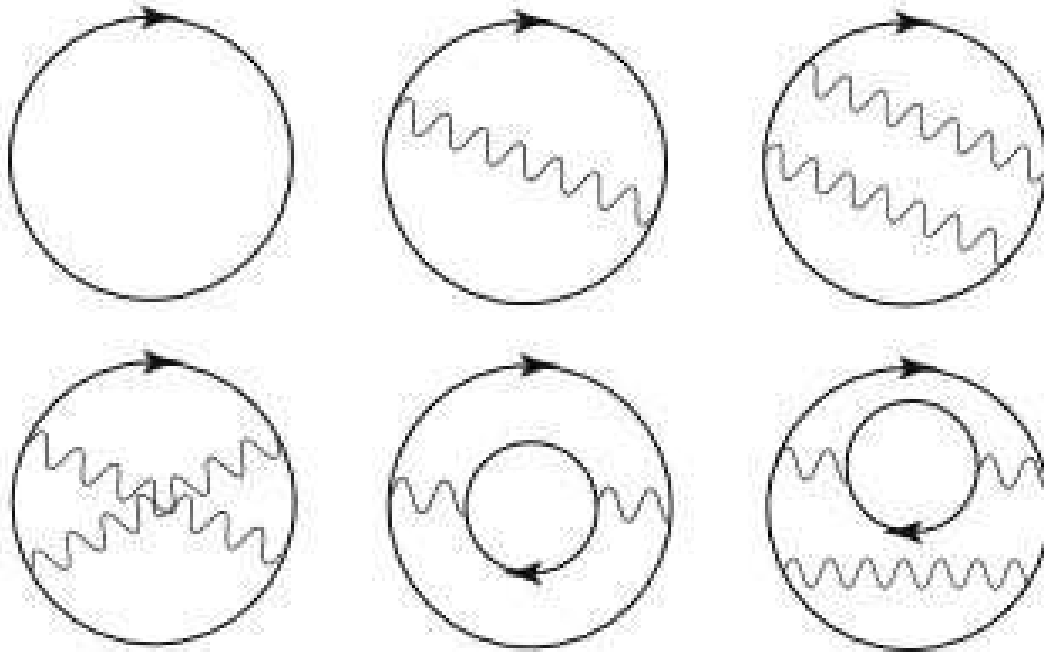
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In modern language the Casimir energy can be expressed in terms of the trace of the Greens function for the fluctuating field in the background of interest (conducting plates)

$$\mathcal{E} = \frac{\hbar}{2\pi} \text{Im} \int d\omega \omega \text{Tr} \int d^3x [\mathcal{G}(x, x, \omega + i\epsilon) - \mathcal{G}_0(x, x, \omega + i\epsilon)]$$

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\mathcal{G} full Greens function for the fluctuating field

\mathcal{G}_0 free Greens function

Trace is over spin

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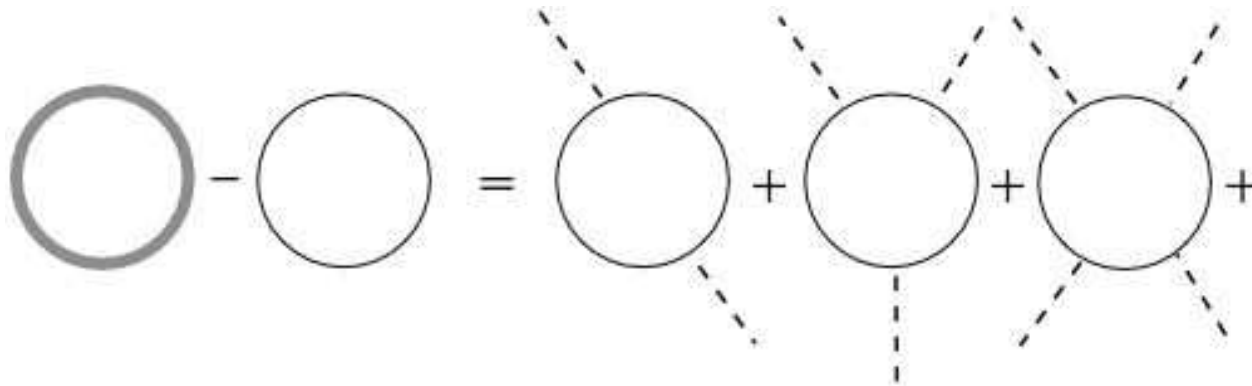
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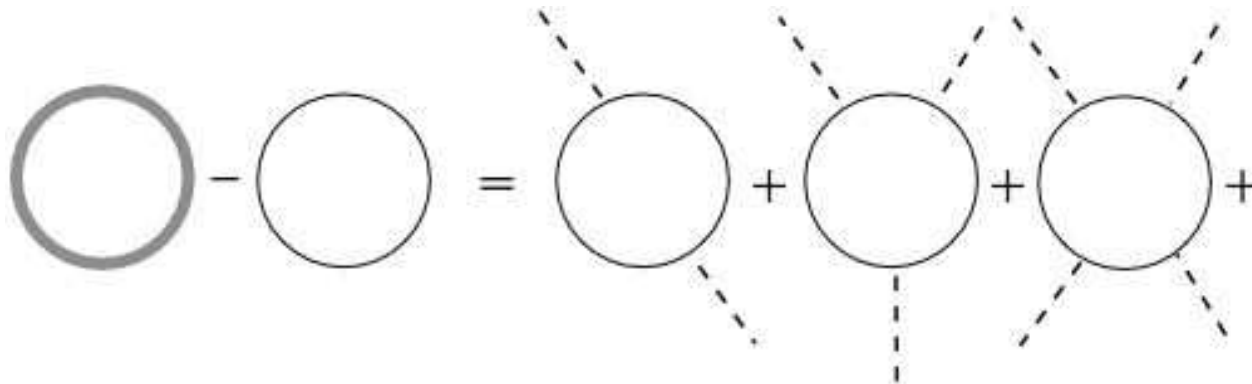
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⇒ “Experimental confirmation of the Casimir effect does not establish the reality of zero point fluctuations” [R. Jaffe et. al.]

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Moore; Razavy, Terning; Johnston, Sarkar; Dodonov et al;
Plunien et al; Barton, Eberlein, Calogeracos; Ford, Vilenkin;
Jaeckel, Reynaud, Lambrecht; Brevik, Milton et al;
Dalvit, Maia-Neto et al; Law; Parentani, ...

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J. Haro & E.E., PRL 97 (2006); arXiv:0705.0597

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- The **dissipative** part we obtain **agrees** with the other methods. But those have problems with the **reactive** part, which in general yields a **non-positive** energy \implies **EXPERIMENT**

SOME DETAILS OF THE METHOD

- **Hamiltonian method** for neutral Klein-Gordon field in a cavity Ω_t , with boundaries moving at a certain speed $v \ll c$, $\epsilon = v/c$ (of order 10^{-8} in [Kim, Brownell, Onofrio, PRL 96 \(2006\) 200402](#))

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- **Hamiltonian.** Transform moving boundary into fixed one by (non-conformal) change of coordinates

$$\mathcal{R} : (\bar{t}, \mathbf{y}) \rightarrow (t(\bar{t}, \mathbf{y}), \mathbf{x}(\bar{t}, \mathbf{y})) = (\bar{t}, \mathbf{R}(\bar{t}, \mathbf{y}))$$

transform Ω_t into a fixed domain $\tilde{\Omega}$

$$\tilde{\Omega}: (t(\bar{t}, \mathbf{y}), \mathbf{x}(\bar{t}, \mathbf{y})) = \mathcal{R}(\bar{t}, \mathbf{y}) = (\bar{t}, \mathbf{R}(\bar{t}, \mathbf{y}))$$

(with \bar{t} the new time)

CASE OF A SINGLE, PARTIALLY TRANSMITTING MIRROR

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Seminal Davis-Fulling model [PRSL A348 (1976) 393]

renormalized energy **negative** while the mirror moves:

cannot be considered as the energy of the produced particles at time t

[cf. paragraph after Eq. (4.5)]

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$$S(\omega) = \begin{pmatrix} s(\omega) & r(\omega) e^{-2i\omega L} \\ r(\omega) e^{2i\omega L} & s(\omega) \end{pmatrix}$$

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\rightarrow Real in the temporal domain: $S(-\omega) = S^*(\omega)$

\rightarrow Causal: $S(\omega)$ is analytic for $\text{Im}(\omega) > 0$

\rightarrow Unitary: $S(\omega)S^\dagger(\omega) = \text{Id}$

\rightarrow The identity at high frequencies: $S(\omega) \rightarrow \text{Id}$, when $|\omega| \rightarrow \infty$

$s(\omega)$ and $r(\omega)$ **meromorphic** (cut-off) functions

(material's **permittivity** and **resistivity**)

RESULTS ARE REWARDING:

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In our Hamiltonian approach

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Note this integral **diverges** for a perfect mirror ($r \equiv -1$, $s \equiv 0$, ideal case), but **nicely converges** for our partially transmitting (physical) one where $r(\omega) \rightarrow 0$, $s(\omega) \rightarrow 1$, as $\omega \rightarrow \infty$

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● The main issue: [S.A. Fulling et. al., hep-th/070209v2](#)

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- **Idea**: zero point fluctuations can contribute to the **cosmological constant** Ya.B. Zeldovich '68

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- Casimir **stress tensor** between pair of parallel perfectly conducting plates, at distance a , transverse dimensions $L \gg a$

[Brown and Maclay, Phys. Rev. 184 (1969) 1272]

$$\langle T_{\mu\nu} \rangle = \frac{\mathcal{E}_c}{a} \text{diag}(1, -1, -1, 3)$$

third spatial direction is normal to plates, \mathcal{E}_c Casimir **energy per unit area**

$$\mathcal{E}_c = -\frac{\pi\hbar c}{720a^3}$$

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- Gravitational interaction of the Casimir apparatus: use gravitational definition of the energy-momentum tensor as **variation of matter part** of action:

$$\delta W_m = \frac{1}{2} \int \sqrt{-g} \delta g_{\mu\nu} T^{\mu\nu} \quad (*)$$

Following **Schwinger** (note the factor 2 in the definition), for a weak field **Fulling et al** define: $g_{\mu\nu} = \eta_{\mu\nu} + 2h_{\mu\nu}$

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energy-momentum tensor of the phys sys must be conserved, so include a
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CC PROBLEM

- Relativistic field: collection of harmonic oscill's (scalar field)

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- What we **do consider** —with relative success in some different approaches— is the **additional** contribution to the cc coming from the **non-trivial topology** of space or from specific **boundary conditions** imposed on braneworld models:

⇒ **kind of cosmological Casimir effect**

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- Assuming one will be able to prove (in the future) that the ground value of the cc is **zero** (as many had suspected until recently), we will be left with this **incremental value** coming from the topology or BCs
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The Braneworld Case

1. Braneworld may help to solve:

- the hierarchy problem
- the cosmological constant problem

2. Presumably, the bulk Casimir effect will play a role in the construction (radion stabilization) of braneworlds

- Bulk Casimir effect (effective potential) for a conformal or massive scalar field
- Bulk is a 5-dim AdS or dS space with 2/1 4-dim dS brane (our universe)
- Consistent with observational data even for relatively large extra dimension

Previous work:

- flat space brane
- bulk conformal scalar field
- conclusion: no CE

We used **zeta regularization** at full power, with **positive** results!

EE, S Nojiri, SD Odintsov, S Ogushi, Phys Rev D67 (2003) 063515 *Casimir effect in de Sitter and Anti-de Sitter braneworlds* EE, SD Odintsov, AA Saharian 0902.0717

Repulsive Casimir effect from extra dimensions and Robin BC: from branes to pistons

Gravity Eqs as Eqs of State: $f(R)$ Case

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- C. Eling, R. Guedens, T. Jacobson [PRL2006]: extension to polynomial $f(R)$ gravity but as non-equilibrium thermodyn.
Also Erik Verlinde (private discussions)

● **Jacobson's argument:** basic thermodynamic relation

$$\delta Q = T\delta S$$

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- **Case of $\mathbf{f}(R)$ gravities:** $\mathbf{L} = \mathbf{f}(R, \nabla^n R)$

- Also the concept of an **effective Newton constant** for graviton exchange (**effective propagator**)

$$\begin{aligned} \frac{1}{8\pi G_{eff}} &= E_R^{pqrs} \epsilon_{pq} \epsilon_{rs} = \frac{\partial \mathbf{f}}{\partial R} (g^{pr} g^{qs} - g^{qr} g^{ps}) \epsilon_{pq} \epsilon_{rs} \\ &= \frac{\partial \mathbf{f}}{\partial R} = \frac{\eta_e}{2\pi}, \quad S = \frac{A}{4 G_{eff}} \end{aligned}$$

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- Final result, for $\mathbf{f}(R)$ gravities:
the local field equations can be thought of as an equation of state of equilibrium thermodynamics (as in the GR case)

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- S-F Wu, G-H Yang, P-M Zhang, arXiv:0805.4044, **direct extension**
 of our results to **Brans-Dicke** and **scalar-tensor** gravities
 T Zhu, Ji-R Ren and S-F Mo, arXiv:0805.1162 [gr-qc];
 C Eling, arXiv:0806.3165 [hep-th]; R-G Cai, L-M Cao and Y-P Hu,
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Moltes gràcies!